



# Demonstration of extreme wavefront sensing and control performance on the TPF High Contrast Imaging Testbed

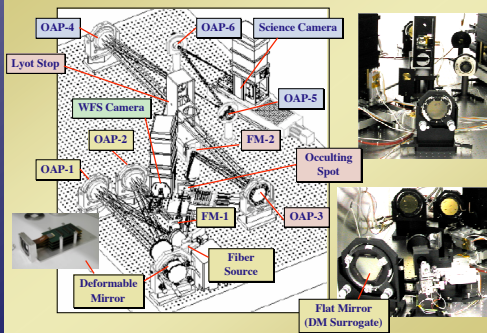
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## Focus Diverse Phase Retrieval

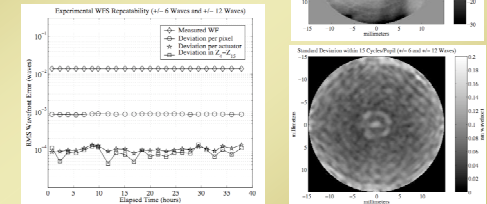
- Image based wavefront sensing (WFS) using focus-diverse phase retrieval
  - Hubble Space Telescope (HST)
  - In flight characterization of primary mirror figure and surface errors
  - Space Infrared Telescope Facility (SIRTF)
  - Used in full system ground testing leading to the accurate focusing of the secondary mirror prior to launch
  - James Webb Space Telescope (JWST)
    - Baseline methodology for finely aligning the segmented primary mirror on orbit
    - In laboratory experiments, it has a demonstrated WFS accuracy  $< \lambda/200$  on small optics and a demonstrated capture range of up to  $10 \lambda$  on large ( $>1$  meter) optics
- Why it is a good choice for a TPF coronagraphic telescope
  - Its large capture range and accuracy makes it an ideal candidate for the initial alignment and correction needed after the telescope deployment
  - The science camera is used for WFS, thus no non-common path optics are employed (other than for pupil-imaging)
- Focus-diverse phase retrieval on the TPF High Contrast Imaging Testbed (HCIT)
  - It has a demonstrated experimental WFS repeatability of  $\lambda/10,000$  (documented in SPIE '03 San Diego paper, results reprinted here)
  - Used to control a  $32 \times 32$  deformable mirror (DM) it has achieved an experimental correction of the sensed wavefront error to the  $\lambda/5,000$  level (presented here)

## High Contrast Imaging Testbed Layout



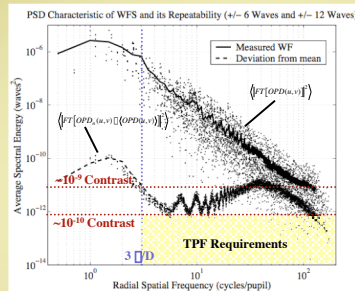
## HCIT WFS Repeatability Experiment

- Experiment conducted before  $32 \times 32$  DM arrivals
  - Surrogate (flat) mirror inserted in its place
- Collected WFS datasets repeatedly for almost 40 hours over a couple of weekdays
  - 26 runs total (24 complete datasets)
- Vacuum tank pumped down to 10 mTorr
  - Pump was on during the experiment
- Temperature remained stable to better than 100 mC



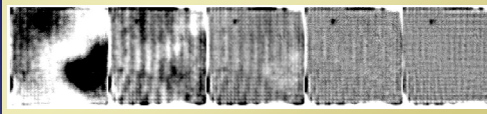
## Power Spectral Density of Experimental WFS Repeatability

- The PSD of the wavefront and its repeatability are frequency distributions of energy
- Increased energy at the lower frequencies are indicative low-order mode variations
- The oscillations are related to the aberration intensity modulation of the defocused PSFs



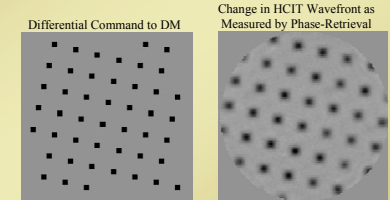
## Deformable Mirror Characterization

- HCIT surface gauge
  - Surface gauge is a Michelson Interferometer
    - Installed on bench in the HCIT vacuum chamber below the HCIT bench
    - In an earlier DM testing effort, it helped to achieve the control a DM to 0.25 Å rms surface error
- Deformable mirror calibration and qualification for HCIT
  - Recently, two Xintetics  $32 \times 32$  DMs were characterized
  - All actuators were commanded to evaluate their gains and influence functions
  - Both DMs were shown to be fully functional and capable of fully correcting themselves along with the surface gauge optical aberrations
  - The initial conservative inter-actuator command constraints prevented the full correction of a single actuator on the second  $32 \times 32$  DM (see bottom figure).



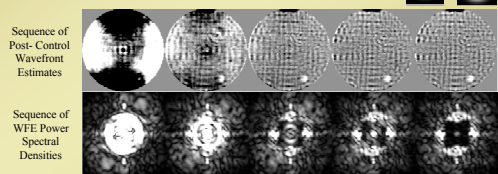
## Calibration of $32 \times 32$ DM for HCIT WFS

- To enable the maximum correction of the HCIT WFE, the DM actuator mapping to the WFS coordinate system must be accurately calibrated
- We applied dilute poke patterns and observed the change in the wavefront using focus diverse phase-retrieval
- Using several such patterns, we calibrated the mapping between the HCIT WFS and the DM coordinate systems and derived the control algorithm



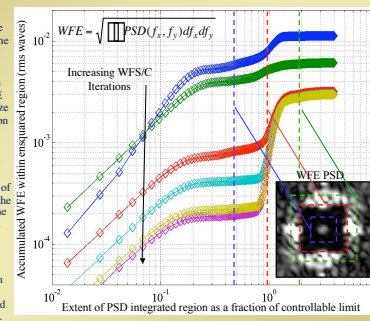
## HCIT Wavefront Sensing and Control

- After the DM was calibrated we conducted an iterative sequence of WFS and WFC.
- With each iteration, improved estimates of the true focus and F/# for HCIT were used to improve the defocus level selection and defocus symmetry in the subsequent focus-diverse data collection
- In this experiment, we have chosen to comply with the conservative inter-actuator position differential constraints. Complying with this operational rule prevents the full correction of 1 actuator within the clear aperture.



## Analysis of WFS/C Performance

- The set of curves depicts the state of the WFE on HCIT over the iterations of WFS/C.
- Each curve represents the accumulated WFE as a function of the size of the ensquared region within the PSD of the wavefront error.
- The WFS/C process achieved a correction of HCIT WFE to about the  $\lambda/5000$  level within the controllable passband
- Beyond this Nyquist limit for the DM, the WFE is a combination of the uncontrollable WFE within HCIT and the DM print-through.



## Initial Contrast Results

- Post-WFS high contrast imaging performance
  - After the last WFS/C iteration, we aligned the coronagraph and measured the system contrast
  - The contrast produced by the corrected system was about  $10^{-6}$  or better over a large portion of the controllable field.
  - This is inconsistent with the wavefront quality implied by the last iteration of WFS/C
  - WFS that was conducted after the coronagraph run indicated that the HCIT wavefront error had increased by 2 nm rms
  - This drift in the wavefront was correlated with temperature variation observed during the experiment
  - This wavefront measurement is consistent with the measurement of the system contrast (See plot)
- Ongoing activities on HCIT
  - Improving HCIT thermal stability
    - The temperature set-point of the chamber has been increased to improve overall stability
  - The DM thermal control system is now active
  - Improving the cycle time to measure contrast
    - This was the first run with the DM and several degrees of freedom had to be aligned within the coronagraph. (4 hours had elapsed)
    - In addition to the Xenon arc-lamp, we are in the process of installing a laser source to reduce our integration times

